

1SWASP J093010.78+533859.5

- 47 Tuc V69
- $\beta$  Aur
- $\chi^2$  Hya
- $\eta$  Mus
- $\psi$  Cen
- $\zeta$  Phe
- AD Boo
- AH Cep
- AI Hya
- AI Phe
- AP And
- AQ Ser
- AR Aur
- ASAS J045304-0700.4
- ASAS J052821+0338.5
- ASAS J082552-1622.8
- ASAS J212954-5620.1
- AY Cam
- BF Dra
- BG Ind
- BK Peg
- BW Aqr
- CD Tau
- CF Tau
- CG Cyg
- CM Dra
- CO And
- CoRoT 102918586
- CoRoT 105906206
- CU Cnc
- CV Boo
- CV Vel
- CW Cep
- DI Her
- DM Vir
- DW Car
- EE Peg
- EF Aqr
- EL Cep
- EK Cep
- EM Car
- EP Cru
- EW Ori
- EY Cep
- FL Lyr
- FS Mon
- FT Ori
- GG Lup
- GG Ori
- GU Boo
- GX Gem
- GZ CMa
- HD 71636
- HP Aur
- HP Dra
- HS Aur
- HS Hya
- HW CMa
- HY Vir
- IM Vir
- IQ Per
- IT Cas
- Kepler-34
- Kepler-35
- KIC 3858884
- KIC 6131659
- KIC 8410637
- KIC 11285625
- KW Hya
- KX Cnc
- LL Aqr
- LSPM J1112+7626
- LV Her
- M4 V65
- M4 V66
- M4 V69
- M55 V54
- MG1-116309
- MG1-2056316
- MG1-506664
- MG1-646680
- MG1-78457
- MU Cas
- MY Cyg

# DEBCat:

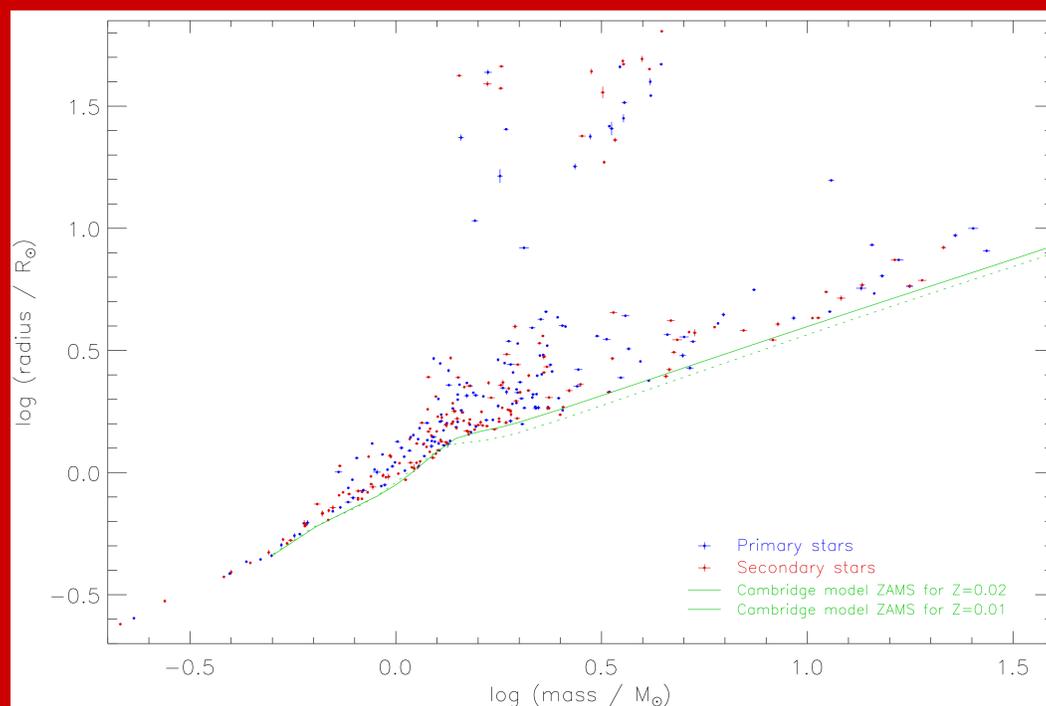
## CATALOGUE OF WELL-STUDIED DETACHED ECLIPSING BINARIES

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Detached eclipsing binary star systems (dEBs) are our primary source of measurements of the physical properties of stars. The masses, radii, and surface gravities of stars in dEBs can be measured empirically and to a precision and accuracy of 1% or better (e.g. Southworth et al., 2005, MNRAS, 363, 529). With an effective temperature measurement we can obtain their luminosities directly. dEBs are the main checks and calibrators for theoretical stellar models (e.g. Pols et al., 1998, MNRAS, 298, 525), and thus form the foundation of stellar and galactic astrophysics.

Important uses of dEBs include as direct distance indicators (Pietrzynski et al. 2013, *Nature*, 495, 76), calibrators of asteroseismic scaling relations (Frandsen et al., 2013, A&A, 556, A138), probes of the chemical evolution of massive stars (Pavlovski et al., 2009, MNRAS, 400, 791), tracers of binary evolutionary processes (Maxted et al., 2013, *Nature*, 498, 463), and characterisation of the host stars of extrasolar planets (Southworth, 2009, MNRAS, 394, 272).

The study of dEBs has a long history (see Stebbins, 1911, ApJ, 34, 112; Russell, 1912, ApJ, 35, 315). Catalogues of well-studied systems have been published by Popper (1967, ARA&A, 5, 85), Popper (1980, ARA&A, 18, 115), Harmanec (1988, BAICz, 39, 329), Andersen (1991, A&ARv, 3, 91) and most recently Torres, Andersen & Giménez (2010, A&ARv, 18, 67).



Plot of the masses and radii of the objects in DEBCat. The zero-age main sequences are from the Cambridge models (Pols et al., 1998, MNRAS, 298, 525) for metal abundances  $Z = 0.02$  and  $0.01$ .

In early 2006 I constructed a catalogue of well-studied dEBs by updating the list of objects compiled by Andersen (1991, A&ARv, 3, 91). From this point onwards new results from refereed journals have been added as and when they are published. The result is DEBCat, which is available at: <http://www.astro.keele.ac.uk/jkt/debcats/>. A changelog has been kept since February 2008. As of September 2014 the total number of objects in DEBCat is 170 (listed on the left and right of this poster).

The requirements for inclusion in DEBCat are:

- ① The evolution of the system has not been significantly affected by binarity. This restricts us to detached binary systems.
- ② The masses and radii of both components have been measured directly, i.e. without significant input from stellar theory, and to an accuracy of 2%. This restricts us to double-lined dEBs. The 2% limit is relaxed for a few interesting objects.
- ③ Measurements of the effective temperatures of both stars are available.

In future I will continue to maintain DEBCat, add measurements of the orbital eccentricity and apsidal motion period, lodge a version with CDS, and redesign the website (currently created from ASCII files by a FORTRAN77 code!).

System	Period (days)	V B-V	Spectral type	Mass (Msun)	Radius (Rsun)	Surface gravity (cgs)	log Teff (K)	log (L/Lsun)	[M/H] (dex)	Ref
V3903 Sgr	1.744	7.27 0.08	O7 V O9 V	27.27 ± 0.55 19.01 ± 0.44	8.088 ± 0.086 6.125 ± 0.060	4.058 ± 0.016 4.143 ± 0.013	4.580 ± 0.021 4.531 ± 0.021	5.087 ± 0.029 4.658 ± 0.032		Vaz et al. (1997A)
V487 Vel	2.753	10.90 0.00	O6 V O6 V	25.3 ± 0.7 8.25 ± 0.17	9.99 ± 0.09 3.49 ± 0.03	3.642 ± 0.018 4.288 ± 0.017	4.559 ± 0.031 4.402 ± 0.046	5.187 ± 0.126 3.649 ± 0.110		Michalska et al.
EM Car	3.414	8.38 -0.31	O8 V O8 V	22.89 ± 0.32 21.43 ± 0.33	9.35 ± 0.17 8.34 ± 0.14	3.856 ± 0.017 3.926 ± 0.016	4.531 ± 0.026 4.531 ± 0.026	5.02 ± 0.10 4.92 ± 0.10		Andersen & Clau
Y Cyg	2.996	7.32 -0.09	O9.8 V O9.8 V	17.72 ± 0.35 17.73 ± 0.30	5.785 ± 0.091 5.816 ± 0.063	4.161 ± 0.014 4.157 ± 0.010	4.521 ± 0.003 4.525 ± 0.003		0.00 ± 0.00	Harmanec et al.
V478 Cyg	2.881	8.63 -0.28	O9.5 V O9.5 V	16.87 ± 0.45 16.31 ± 0.35	7.423 ± 0.079 7.423 ± 0.079	3.919 ± 0.015 3.909 ± 0.013	4.484 ± 0.015 4.485 ± 0.015	4.63 ± 0.06 4.63 ± 0.06		Popper & Hill (19 Popper & Etzel (1 Holmgren et al.
AH Cep	1.775	6.81 0.30	B0.5 V B0.5 V	15.2 ± 0.2 13.6 ± 0.2	6.38 ± 0.11 5.86 ± 0.13	4.02 ± 0.04 4.04 ± 0.05	4.476 ± 0.029 4.456 ± 0.015			
V578 Mon	2.408	8.54 0.17	B1 V early B	14.54 ± 0.08 10.29 ± 0.06	5.41 ± 0.04 4.20 ± 0.05	4.133 ± 0.018 4.185 ± 0.021	4.477 ± 0.007 4.411 ± 0.007	4.33 ± 0.03 3.86 ± 0.03	0.00 ± 0.00	Garcia et al. (ar Pavlovski & Hen
V453 Cyg	3.890	8.29 0.18	B0.4 IV B0.7 IV	14.36 ± 0.20 11.11 ± 0.13	8.551 ± 0.055 5.489 ± 0.063	3.731 ± 0.012 4.005 ± 0.015	4.446 ± 0.006 4.410 ± 0.008	4.593 ± 0.025 4.098 ± 0.034	0.00 ± 0.00	Southworth et a Pavlovski & Sout
CW Cep	2.729	7.59 -0.28	B0.5 V B0.5 V	13.52 ± 0.39 12.08 ± 0.29	5.685 ± 0.130 5.177 ± 0.129	4.059 ± 0.024 4.082 ± 0.024	4.452 ± 0.016 4.442 ± 0.016	4.27 ± 0.06 4.15 ± 0.07		Clausen & Cime Andersen (1991)
V380 Cyg	12.426	5.68 -0.06	B1.5 III B2 V	11.43 ± 0.19 7.00 ± 0.14	15.71 ± 0.13 3.819 ± 0.048	3.104 ± 0.006 4.120 ± 0.011	4.336 ± 0.006 4.356 ± 0.023	4.691 ± 0.041 3.626 ± 0.038	0.03 ± 0.10	Tkachenko et al Tkachenko et al.
DW Car	1.328	9.68 0.07	B1 V B1 V	11.34 ± 0.12 10.63 ± 0.14	4.558 ± 0.045 4.297 ± 0.055	4.175 ± 0.008 4.198 ± 0.011	4.446 ± 0.016 4.423 ± 0.016	4.055 ± 0.063 3.915 ± 0.067	0.00 ± 0.00	Southworth & C Clausen et al. (2

- NGC 7142 V1
- OGLE-LMC-CEP0227
- OGLE-LMC-ECL01866
- OGLE-LMC-ECL03160
- OGLE-LMC-ECL06575
- OGLE-LMC-ECL09114
- OGLE-LMC-ECL09660
- OGLE-LMC-ECL10567
- OGLE-LMC-ECL15260
- OGLE-LMC-ECL26122
- OGLE SMC126.1 210
- OGLE SMC113.3 4007
- OGLE SMC130.5 4296
- OGLE SMC101.8 14077
- OGLE SMC108.1 14904
- PT Vel
- PV Cas
- PV Pup
- QX Car
- RR Lyn
- RS Cha
- RT CrB
- RW Lac
- RZ Cha
- SW CMa
- SZ Cen
- TV Nor
- TZ For
- TZ Men
- U Oph
- UX Men
- UZ Dra
- V335 Ser
- V364 Lac
- V375 Cep
- V380 Cyg
- V392 Car
- V396 Cas
- V404 CMa
- V413 Ser
- V432 Aur
- V442 Cyg
- V451 Oph
- V453 Cyg
- V459 Cas
- V467 Vel
- V478 Cyg
- V501 Her
- V539 Ara
- V565 Lyr
- V568 Lyr
- V578 Mon
- V596 Pup
- V624 Her
- V636 Cen
- V760 Sco
- V785 Cep
- V885 Cyg
- V906 Sco
- V1031 Ori
- V1061 Cyg
- V1130 Tau
- V1143 Cyg
- V1174 Ori
- V1229 Tau
- V1236 Tau
- V1388 Ori
- V1647 Sgr
- V2080 Cyg
- V2365 Oph
- V3903 Sgr
- V4089 Sgr
- VV Crv
- VZ Cep
- VZ Hya
- WOCS 40007
- WW Aur
- WW Cam
- WX Cep
- WZ Oph
- XY Cet
- Y Cyg
- YY Gem
- YZ Cas
- ZZ UMa

<http://www.astro.keele.ac.uk/jkt/debcats/>