

Spitzer Studies of Core-collapse supernovae

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+ Mid-IR Supernova Consortium

Imperial College London (Meikle, Gerardy)

Cornell Uni. (Farrah)

Dark Cosmology Centre, Copenhagen (Andersen)

Florida State Uni. (Hoflich)

Queen's Uni. Belfast (Mattila)

Spitzer Science Center (van Dyk)

Stockholm Observatory (Fransson, Lundqvist, Sollerman)

Uni. of California, Berkeley (Filippenko, Foley)

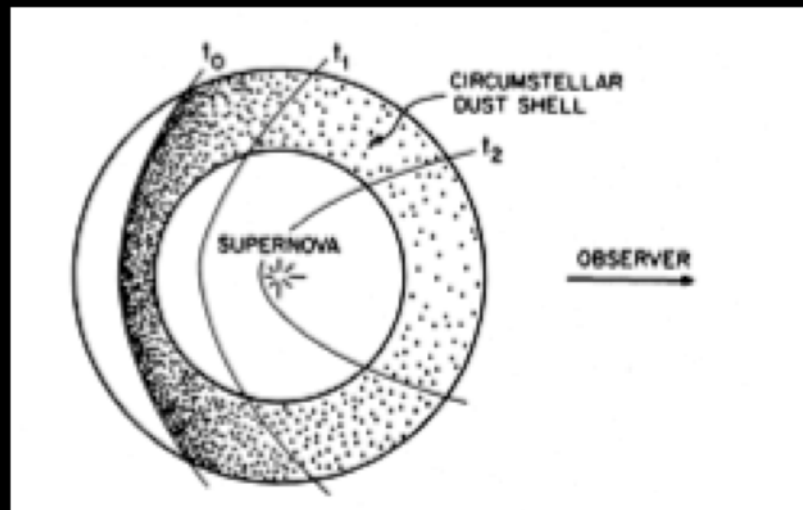
Uni. Of Texas, Austin (Wheeler)

- When does dust form?
- Under what conditions? → Molecules always necessary?
- How much dust?
- What is the composition?

Thermal emission

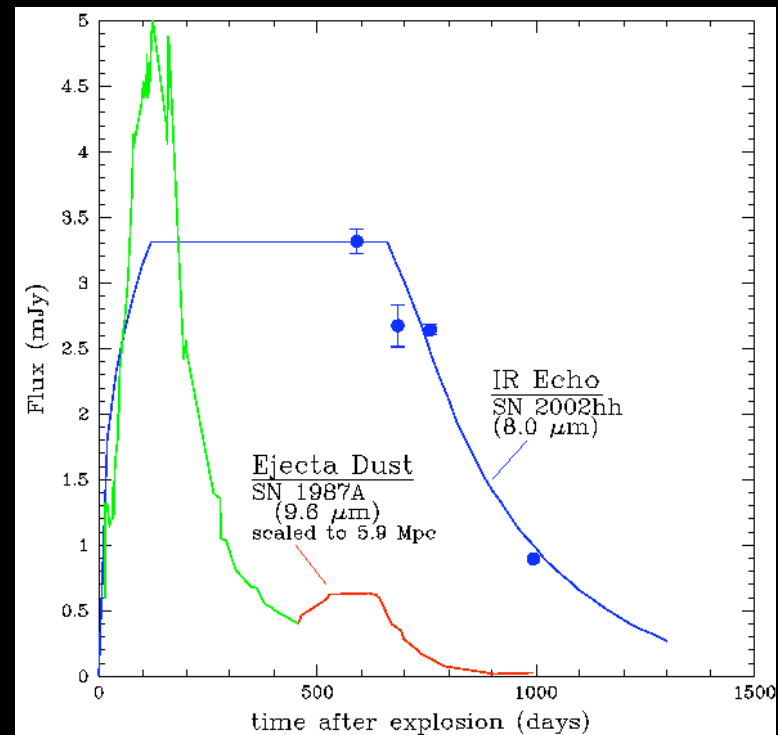
Direct detection of dust

..... But could be due to pre-existing dust in the circumstellar medium (CSM)



Dwek (1983)

⇒ Monitoring



Bouchet et al. 1989, 93
Meikle et al 2006

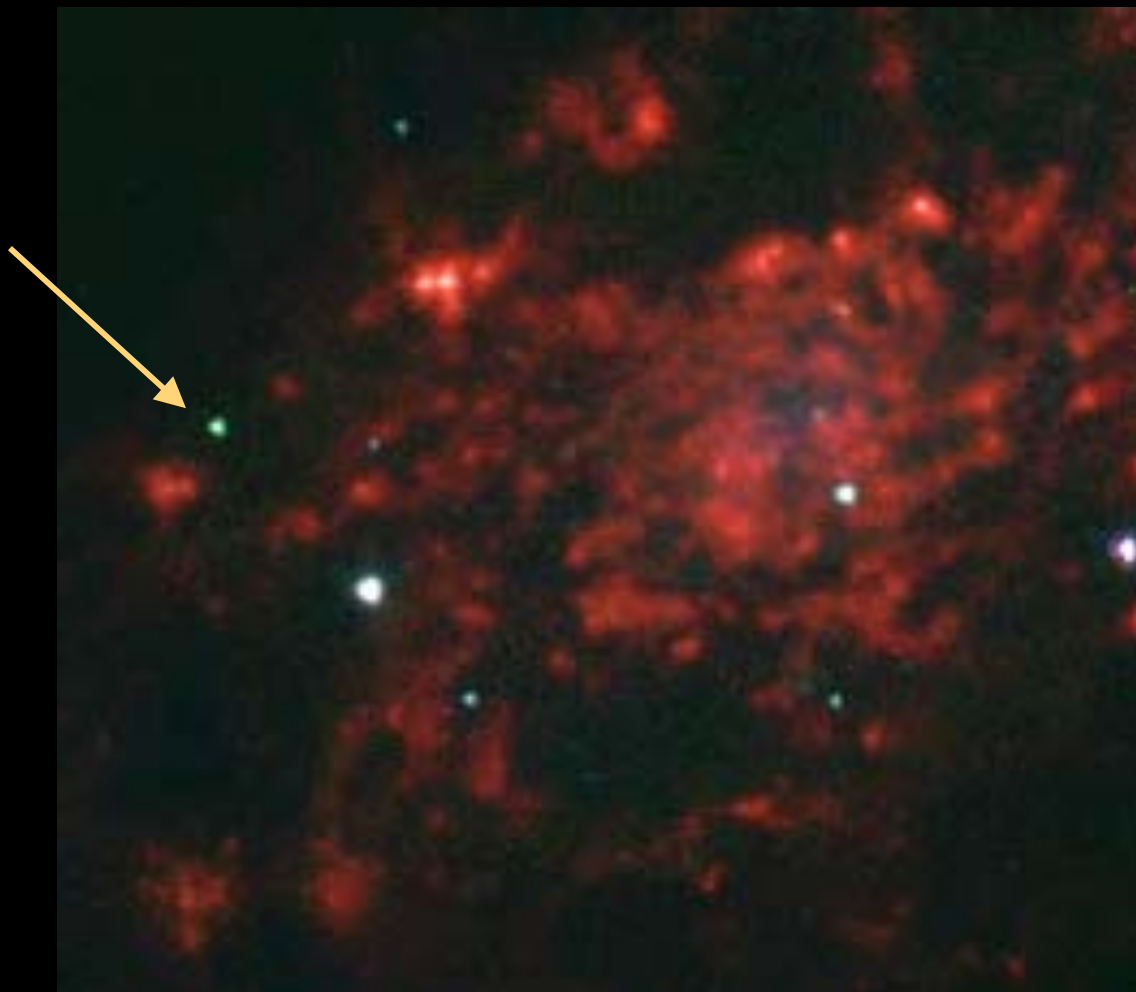
Mid-IR monitoring extremely challenging from the ground.

No mid-IR studies of SNe since SN 1987A

MISC Supernova programme with Spitzer:

- ~145 hrs over cycles 1,2,3
- Sample of nearby targets (photometry + spectroscopy)
- to assess the ubiquity of dust formation in core-collapse supernovae

SN 2004dj (type II-P)



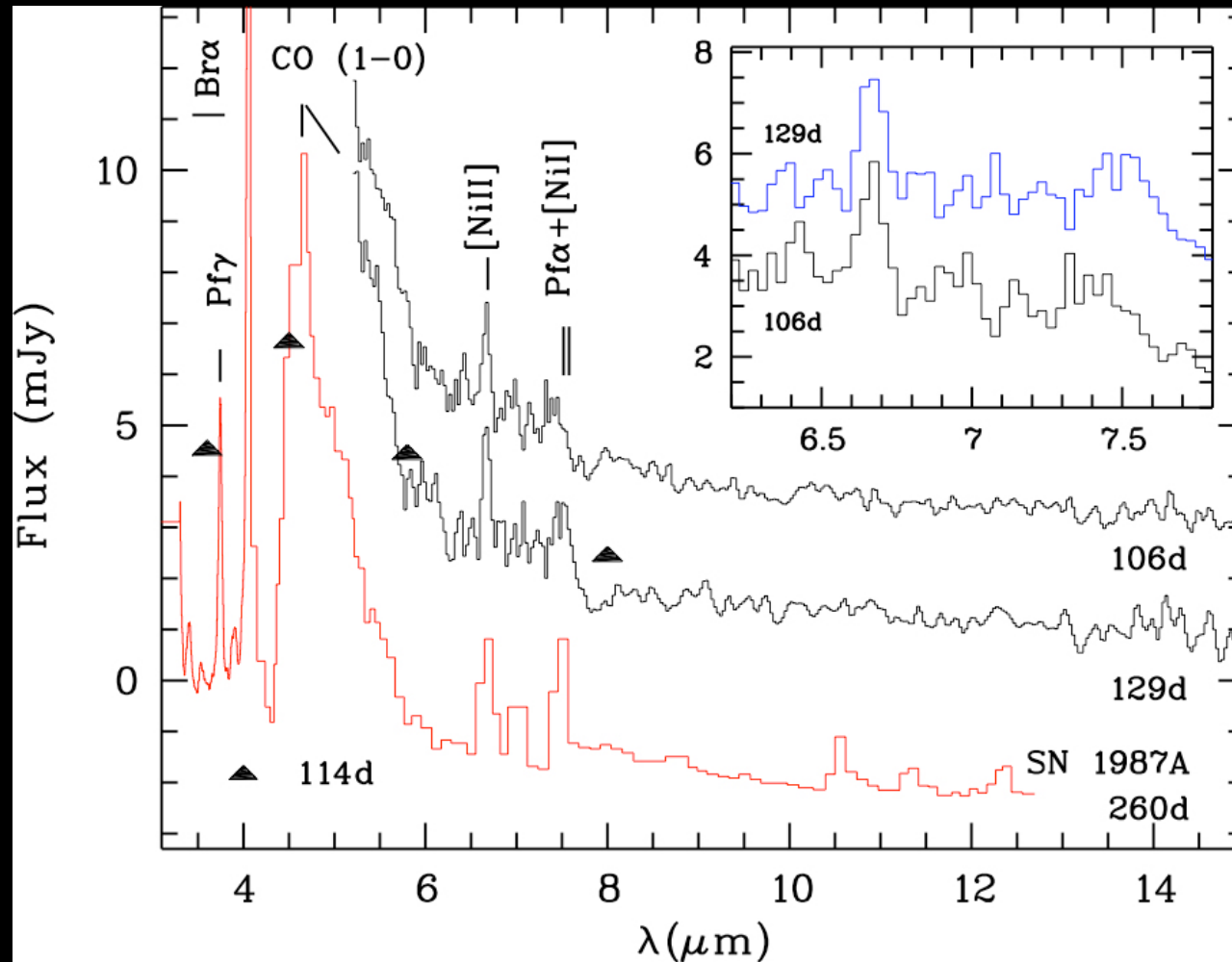
Blue 3.6 μ m

Green 4.5 μ m

Red 8.0 μ m

The first mid-IR spectrum of any SN since SN 1987A

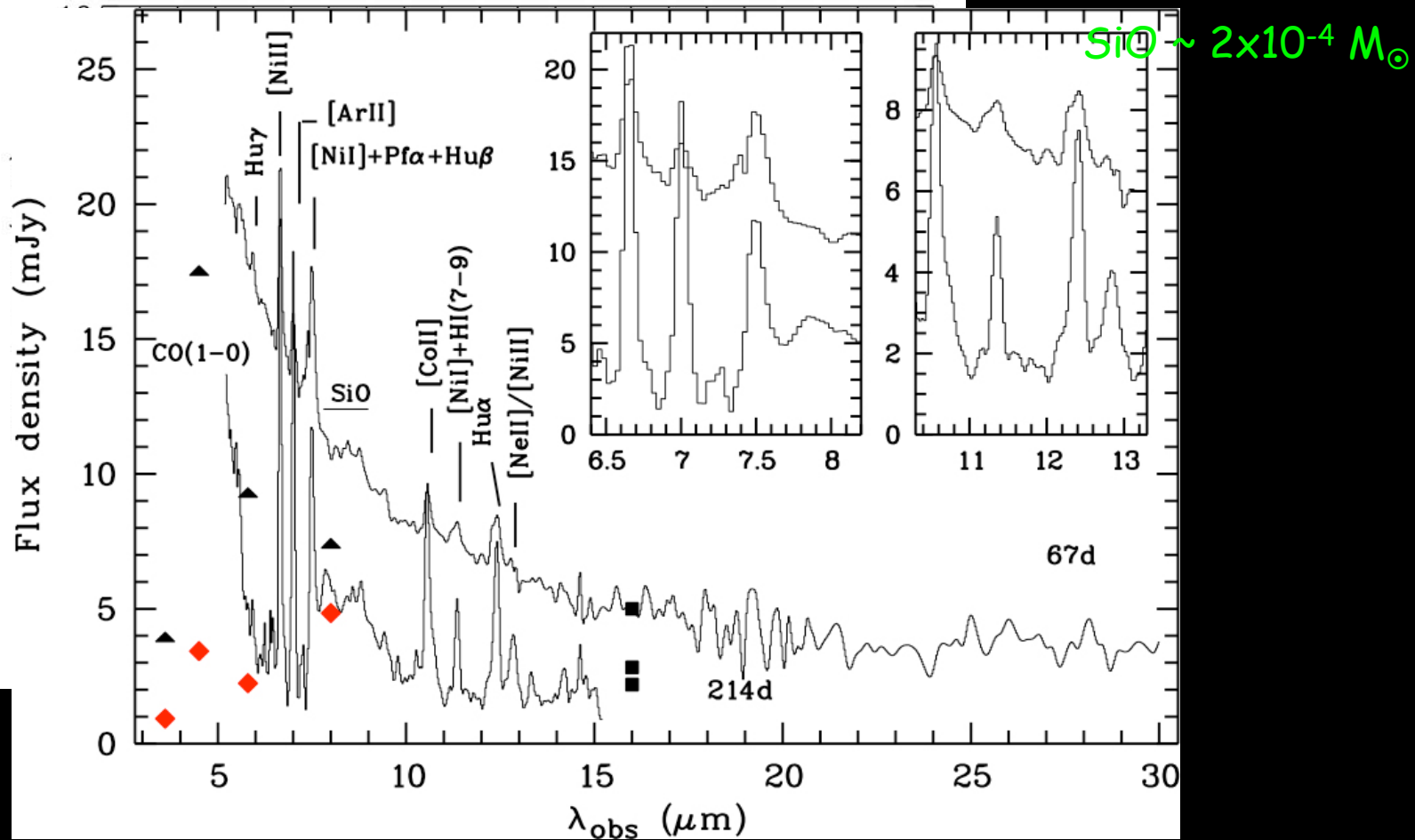
SN 2004dj



Kotak et al. (2005); Wooden et al. (1991)

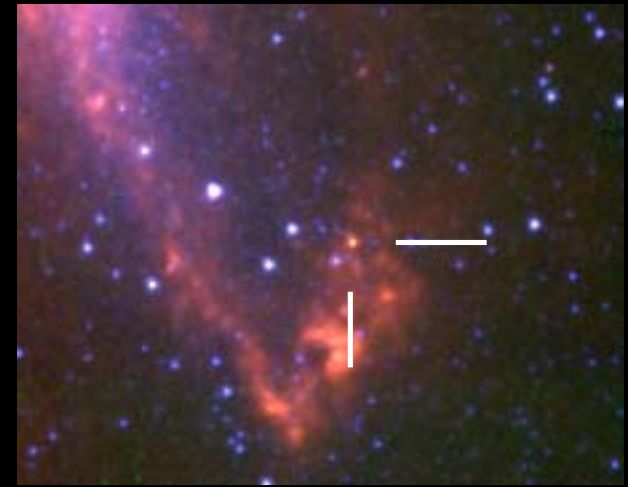
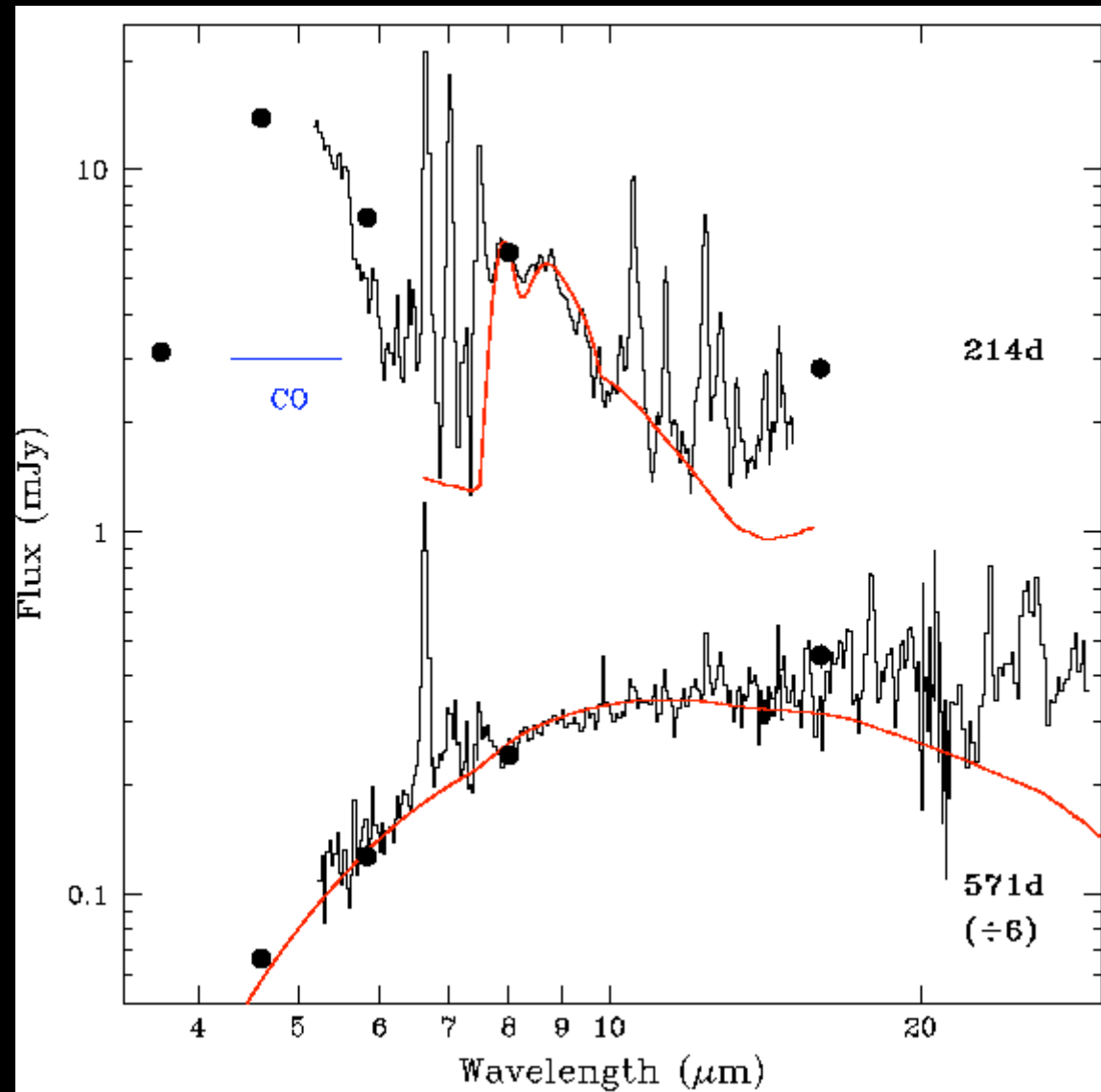
First detection of SiO since SN 1987A

SN 2005af (type IIP)



Kotak et al. (2006)

SN 2005af



Spitzer IRAC 3.6-8 μm 576d

Age (d)	Temp. (K)	Md ($10^{-4} M_{\odot}$)	f
214	800	0.015	0.07
571	420	4.0	0.32

Recent claims of a large dust mass in SN 2003gd (IIP)

0.02 M_{\odot} of dust -- Sugerman et al. (Science, 2006)

Sugerman et al. (2006) Meikle et al. (2007)

24 μ m / ~700d 106 ± 16 380 ± 90 μ Jy (same data)

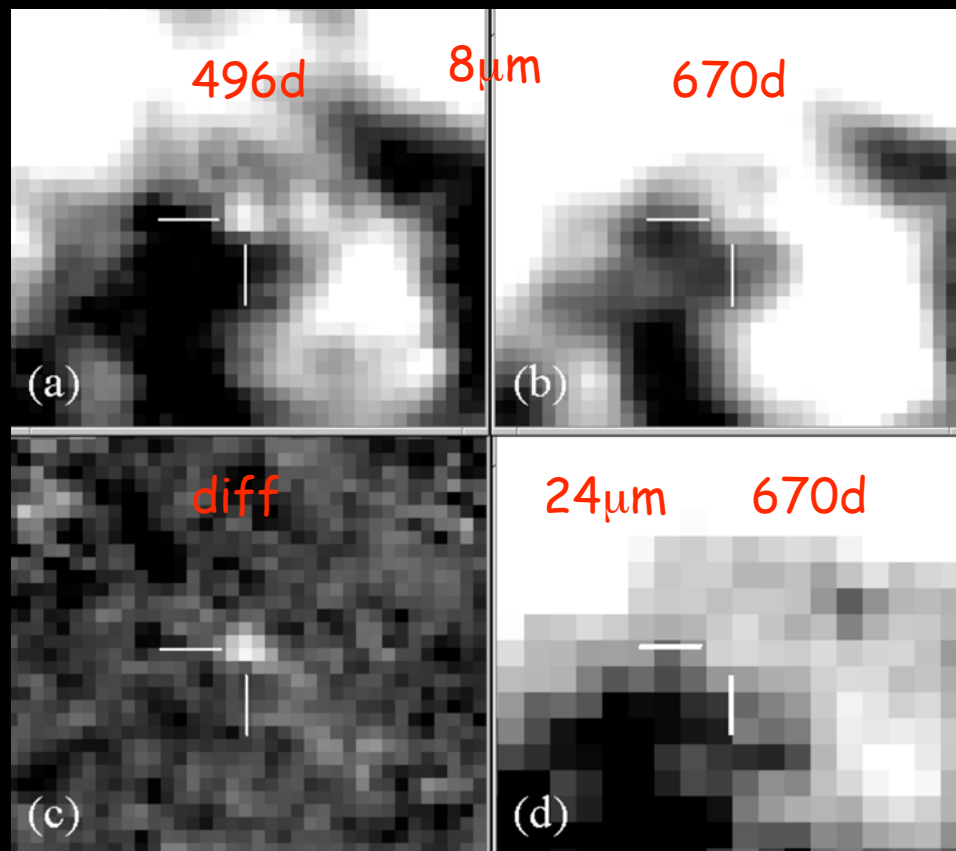
Sensitivity overestimate: $106 \mu\text{Jy} \Rightarrow S/N \sim 1$ (not 6.5)

Error in 24 μ m / ~700d flux of Sugerman et al.

- Outer limit of dust-forming zone > metal line velocities from late-time spectra (~2000km/s) \Rightarrow unphysical
- Total luminosity > 4 x total radioactive luminosity deposited in ejecta
 \Rightarrow severe energy deficit

Directly detected dust: SN 2003gd produced no more than few $10^{-5} M_{\odot}$

For details see Meikle et al. (2007)



Summary

- CO and SiO detected in all of our nearby type IIP sample
- Before Spitzer, no evidence of dust formation in the most common type of core-collapse SNe (IIP).
- We find evidence for (direct) dust condensation in all of our type IIP events.
- Currently only $10^{-3} - 10^{-5} M_{\odot}$ of dust 10-100x lower than needed.
-- more may exist in optically-thick clumps

Core-collapse supernovae as massive dust factories?

"Definitely Maybe"